# New Vacuum Solutions for Quadratic Metric–affine Gravity

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University College London 28 August 2008

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- PP-waves with torsion

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- Interpretation

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 Spacetime considered to be a connected real 4–manifold M equipped with a Lorentzian metric g and an affine connection Γ, i.e.

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- **Definition.** We call a spacetime  $\{M, g, \Gamma\}$  *Riemannian* if the connection is Levi–Civita (i.e.  $\Gamma^{\lambda}_{\mu\nu} = \left\{ \begin{matrix} \lambda \\ \mu\nu \end{matrix} \right\}$ ), and *non-Riemannian* otherwise.

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Action is

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• The quadratic form *q*(*R*) has 16 *R*<sup>2</sup> terms with 16 real coupling constants.

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- The quadratic form *q*(*R*) has 16 *R*<sup>2</sup> terms with 16 real coupling constants.
- Action conformally invariant, unlike Einstein-Hilbert.

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 Independent variation of g and Γ produces the system of Euler–Lagrange equations

$$\partial S/\partial g = 0,$$
 (1)  
 $\partial S/\partial \Gamma = 0.$  (2)

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- Einstein spaces (Yang, Mielke);
- pp-waves with parallel Ricci curvature (Vassiliev);
- Certain explicitly given torsion waves (Singh and Griffiths);
- Triplet ansatz (Hehl, Macías, Obukhov, Esser, ...);
- Minimal pseudoinstanton generalisation (Obukhov).

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- **Definition.** A *pp-wave* is a Riemannian spacetime whose metric can be written locally in the form

$$\mathrm{d}s^2 = 2\,\mathrm{d}x^0\,\mathrm{d}x^3 - (\mathrm{d}x^1)^2 - (\mathrm{d}x^2)^2 + f(x^1,x^2,x^3)\,(\mathrm{d}x^3)^2$$

in some local coordinates.

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in some local coordinates.

 Well known spacetimes in GR, simple formula for curvature - only trace free Ricci and Weyl pieces.

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Plane wave solutions of this equation can be written down as

$$A = h(\varphi) m + k(\varphi) I,$$
  
$$\varphi : M \to \mathbb{R}, \qquad \varphi(x) := \int I \cdot dx.$$

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• **Definition** A *generalised pp-wave* is a metric compatible spacetime with pp-metric and torsion

$$T:=\frac{1}{2}\operatorname{Re}(A\otimes dA).$$

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• Curvature of a generalised pp-wave is

$$R = -\frac{1}{2}(I \wedge \{\nabla\}) \otimes (I \wedge \{\nabla\})f + \frac{1}{4}\operatorname{Re}\left((h^2)''(I \wedge m) \otimes (I \wedge m)\right).$$

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• Torsion of a generalised pp-wave is

$$T = \operatorname{Re}\left((a \ l + b \ m) \otimes (l \wedge m)\right),$$

where

$$a := \frac{1}{2} h'(\varphi) k(\varphi), \quad b := \frac{1}{2} h'(\varphi) h(\varphi).$$

#### Main result of the thesis

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- In special local coordinates, 'parallel Ricci curvature' is written as  $f_{11} + f_{22} = \text{const.}$

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- **Theorem** Generalised pp-waves of parallel Ricci curvature are solutions of the field equations (1) and (2).
- In special local coordinates, 'parallel Ricci curvature' is written as  $f_{11} + f_{22} = \text{const.}$
- Generalised pp-waves of parallel Ricci curvature admit a simple explicit description.

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- We write down the field equations (1) and (2) for general metric compatible spacetimes and substitute the formulae for torsion, Ricci and Weyl into these.
- Together with  $\nabla Ric = 0$ , we get the result.
- This result was first presented in : "PP-waves with torsion and metric affine gravity", 2005 V. Pasic, D. Vassiliev, *Class. Quantum Grav. 22 3961-3975*.

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- Torsion and torsion generated curvature are waves traveling at the speed of light.
- Underlying pp-space can be viewed as the 'gravitational imprint' created by wave of some massless field.
- Mathematical model for some massless particle?

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Look at Weyl action

$$S_{W} := 2i \int \left( \xi^{a} \sigma^{\mu}_{a\dot{b}} \left( \nabla_{\mu} \bar{\xi}^{\dot{b}} \right) - \left( \nabla_{\mu} \xi^{a} \right) \sigma^{\mu}_{a\dot{b}} \bar{\xi}^{\dot{b}} \right),$$

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In a generalised pp-space Weyl's equation takes form

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In a generalised pp-space Weyl's equation takes form

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• There exist pp-wave type solutions of Einstein-Weyl model

$$egin{aligned} S_{\mathrm{EW}} &:= k \int \mathcal{R} + S_{\mathrm{neutrino}}, \ \partial S_{\mathrm{EW}} / \partial g &= 0, \ \partial S_{\mathrm{EW}} / \partial \xi &= 0. \end{aligned}$$